

LARGE Antennas in Space: Concepts, Options and Challenges

Yahya Rahmat-Samii
Department of Electrical Engineering
UCLA

Los Angeles, CA 90095-1594
email: rahmat@ee.ucla.edu
http: //www.ee.ucla.edu
Tel: (310) 206 3847; Fax: (310) 206 8495

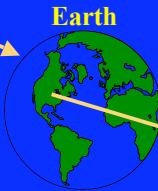
April 14, 2004



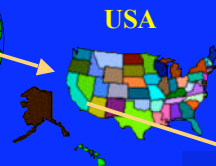
UCLA: University of California, Los Angeles



Solar system



Earth



USA

Los Angeles, CA



Santa Monica
Pacific Ocean

Bel Air

UCLA
Campus

Beverly
Hills

Westwood Village



EE
Department



	EE Department:2002-03 UCLA Faculty: 47 Lecturers: 17 Admin. Staff: 48	
Publications Paper: 139 Conference: 264 Book/Chapter: 14 Patent Issued: 7 Many Awards		Research Fund: \$22.8 M Supports: Fellowships RA-Ships TA-Ships Readers
	Students Undergrad Students: 752 Graduate Students: 444 Degrees Conferred: B.S.: 172, M.S.: 111, Ph.D.: 29	

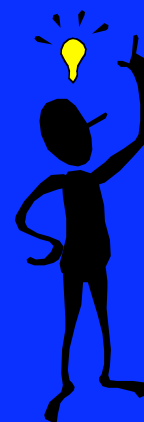
Universities must be considered as vital players in creating new and visionary concepts for future missions.

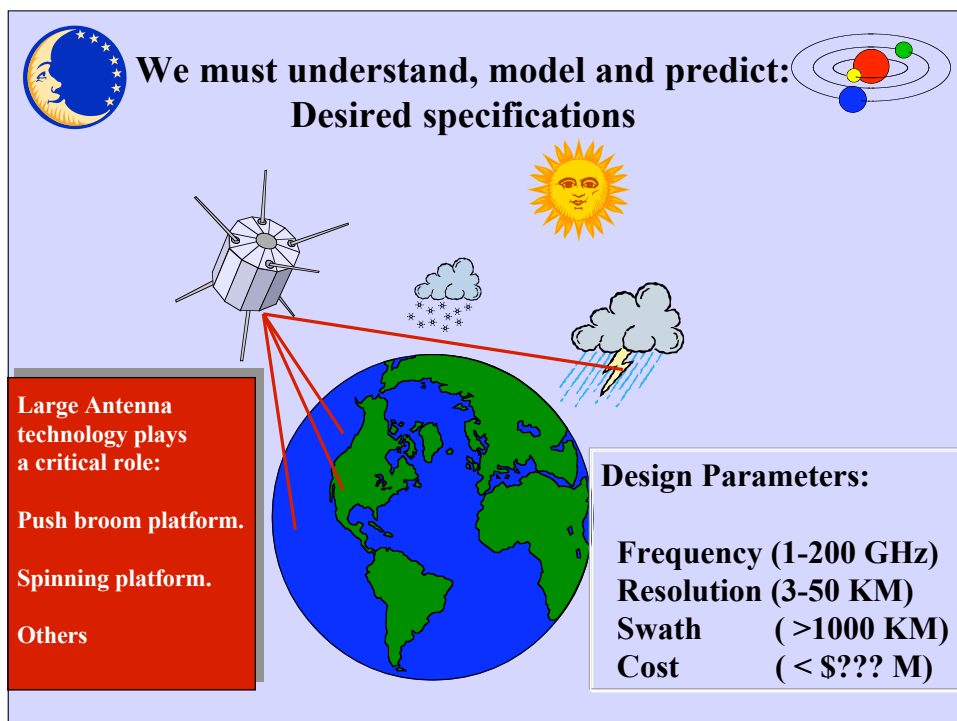
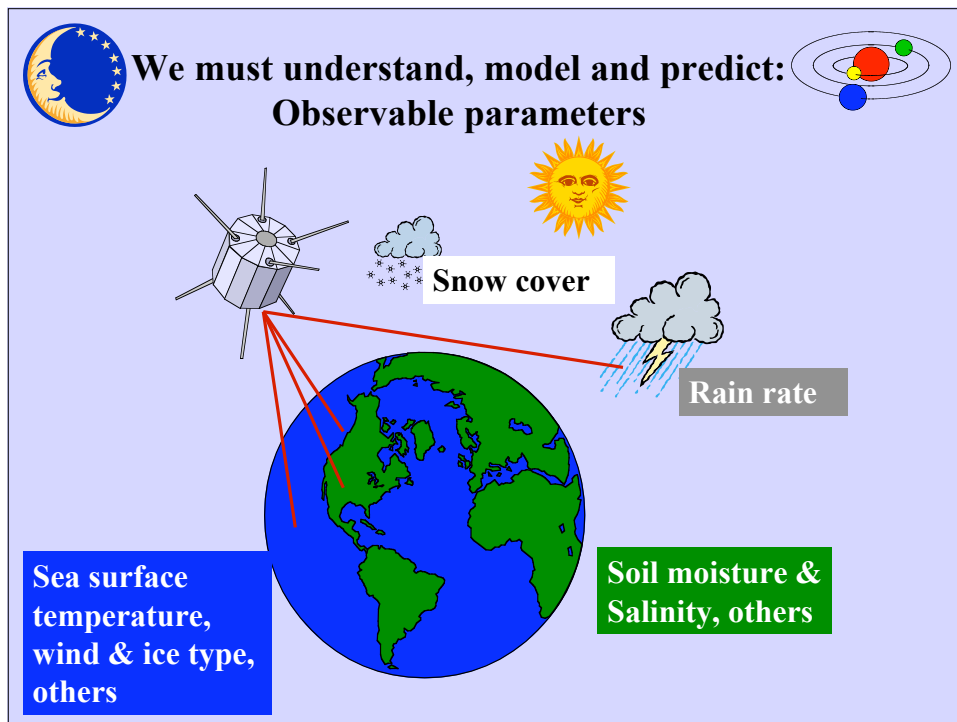


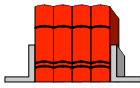
One needs to dream big!

**"It is difficult to say what is impossible,
for the dream of yesterday is the hope of
today and the reality of tomorrow."**

R. H. Goddard







How does an antenna radiate?

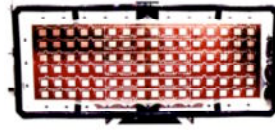
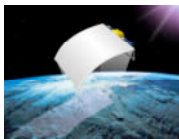
Fact from physics: Accelerated charges radiate.

$I = q v$
current = charge x velocity

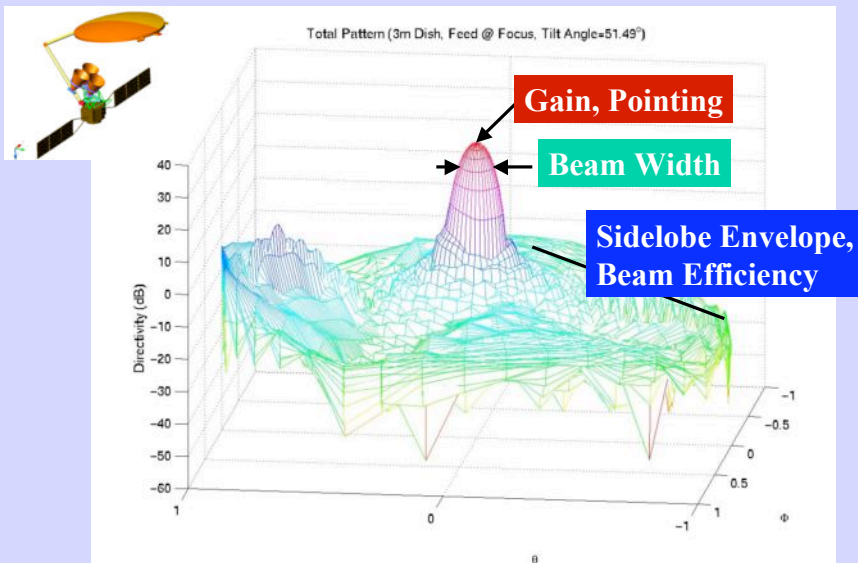
$dI/dt = q dv/dt$
time varying
current = accelerated charge

Observation: Time varying currents produce electromagnetic radiation governed by Maxwell's equations.

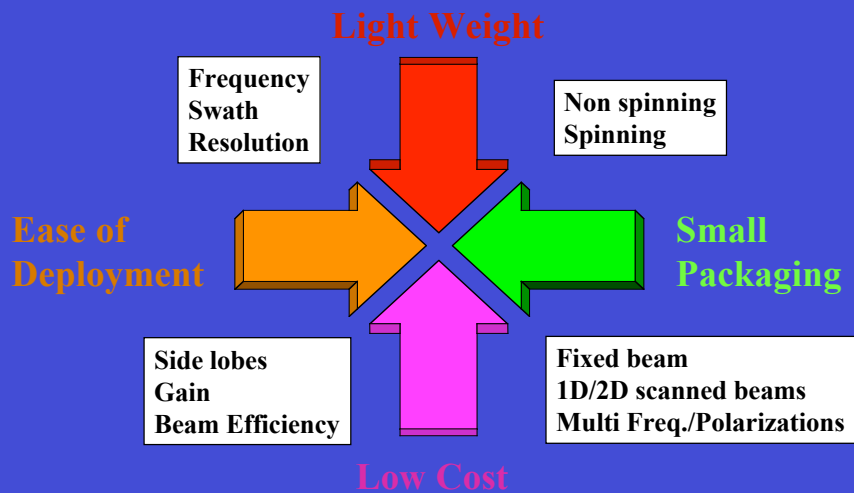
Antennas: Electromagnetic devices controlling the flow of the time varying currents; thus, producing EM radiation with desired characteristics.



Antenna pattern is the manifest of the spatial Fourier transform of its current distribution

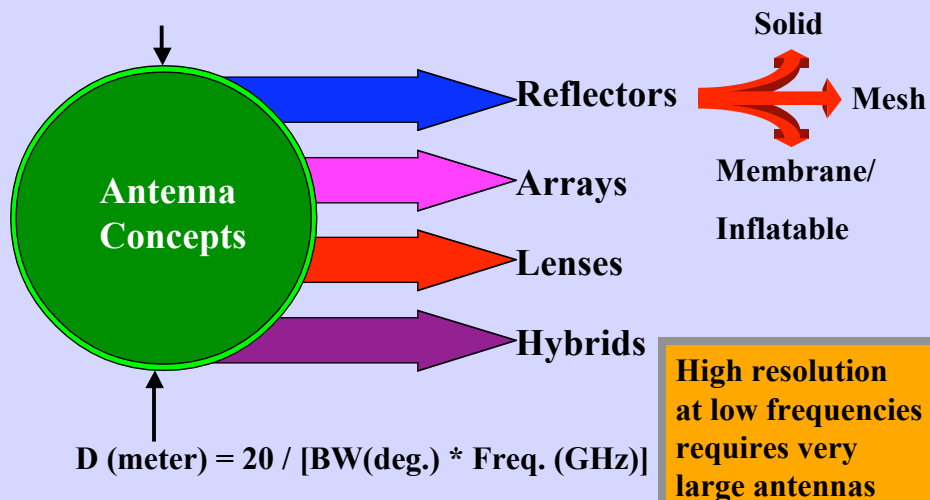


System Parameters for Large Space Antennas



The challenge is how to achieve required performance!

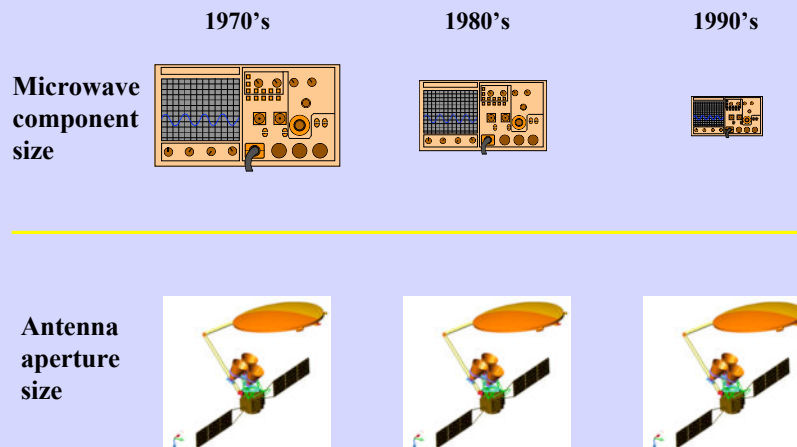
Antenna Classifications



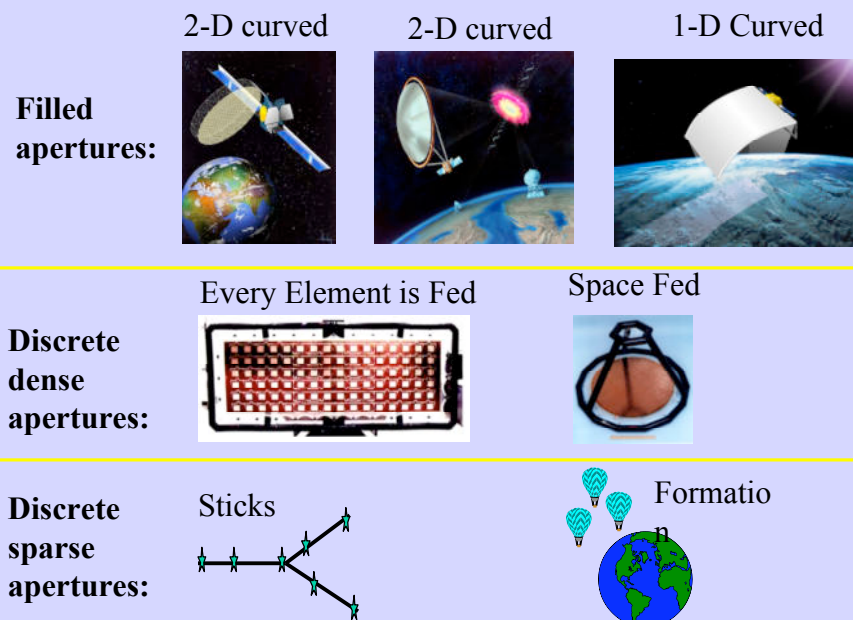
$$D \text{ (meter)} = 20 / [\text{BW(deg.)} * \text{Freq. (GHz)}]$$

Example: $D \text{ (meter)} = 20 / [1.1(\text{deg.}) * 1.4 \text{ (GHz)}] = 13.0 \text{ meters}$

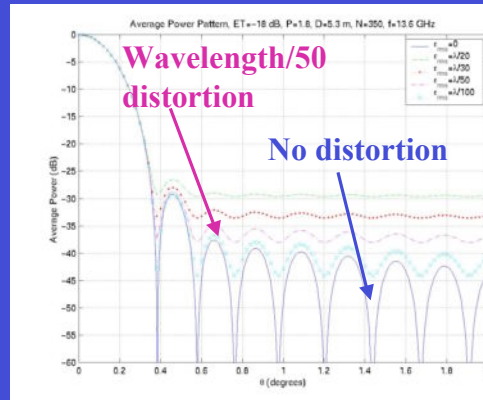
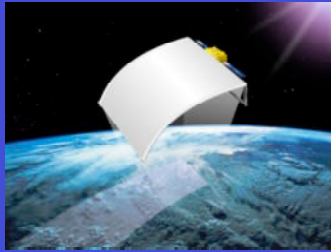
Microwave Components are becoming smaller!
Antenna aperture stays LARGE for desired resolution!



How to obtain LARGE apertures?



Effects of surface RMS on antenna pattern



Depending on the desired sidelobe envelope, surface RMS better than wavelength/50 (for reflector antennas) and wavelength/25 (for array antennas) may be required.



Reflector Design Options



Symmetric Parabolic

Offset Parabolic

Spherical

Cylindrical

Torus

Cassegrain

Gregorian

Others



Feed Design Options



Conical Horns

Corrugated Horns

Patch arrays

Multi
Polarizations

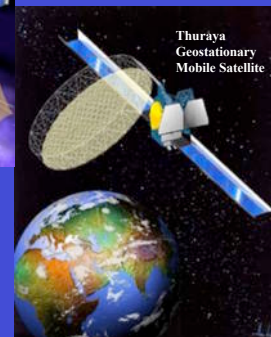
Multi
Frequencies

Others

Satellite Communication Industry is relying on large deployable mesh reflector antennas

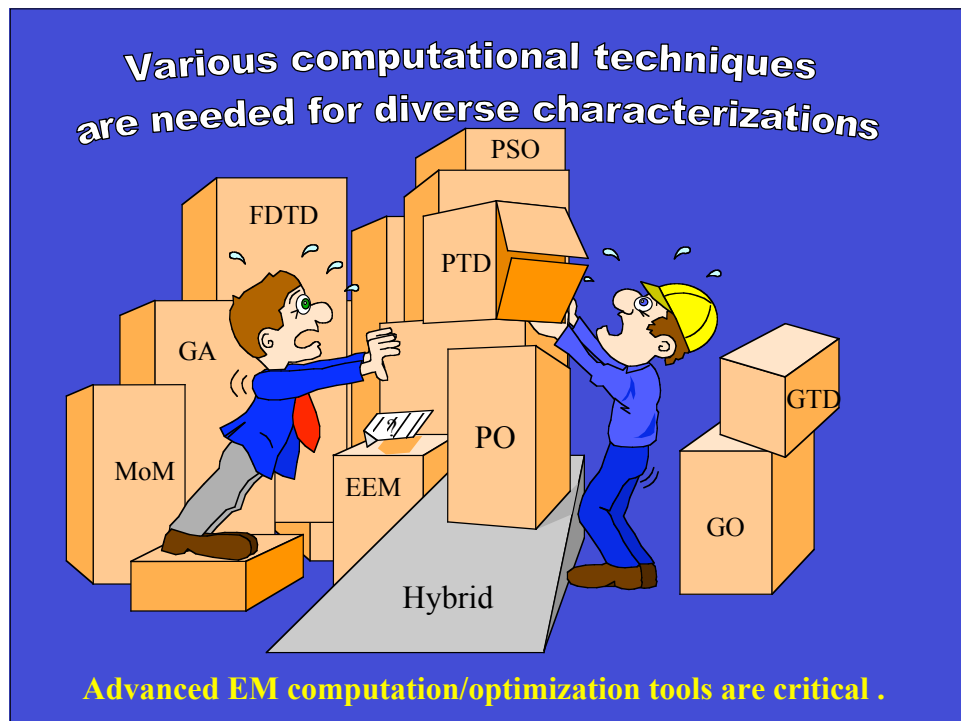


12 m antenna
2000-2004

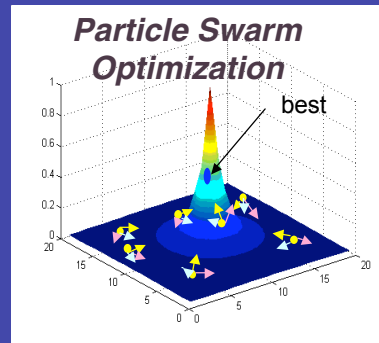


A. G. Roederer and Y. Rahmat-Samii, "Unfurable Satellite Antennas: A Review," Annales des Telecommunications, vol. 44, no. 9-10, pp. 475-488, November, 1989.

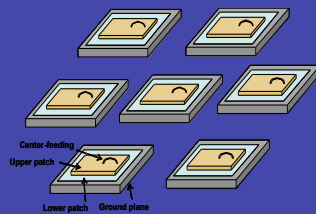
Electromagnetic Computations



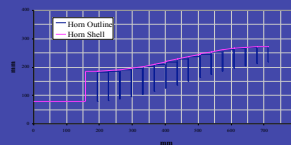
Application Examples: Potential Feeds for Reflector Antennas



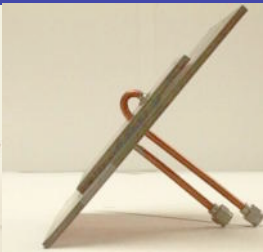
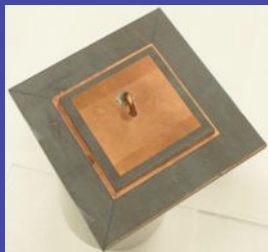
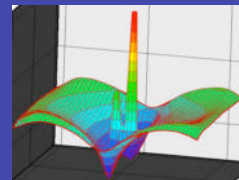
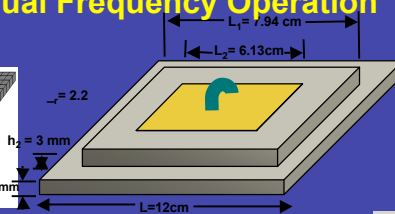
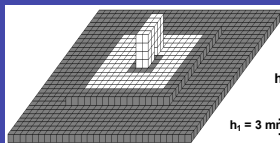
Array of Stacked Patch antennas



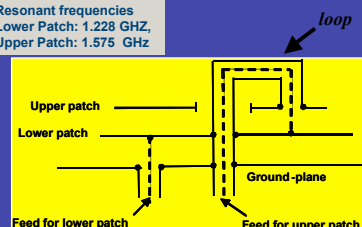
Optimized Profiled horn



Prototype of the Center-Fed Stacked Patch Dual Frequency Operation



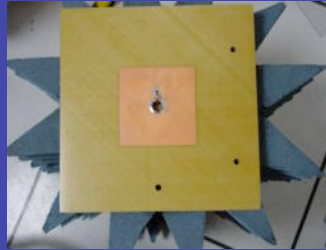
Resonant frequencies
Lower Patch: 1.228 GHz,
Upper Patch: 1.575 GHz



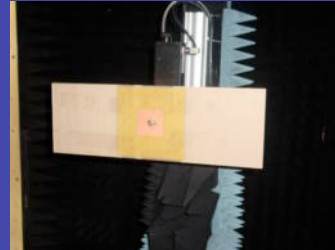
Side-view of the stacked patch

Test model of the Center-fed Stacked Patch built at UCLA

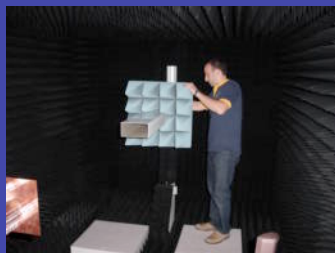
Stacked Patch in the Spherical Near Field Measurement Chamber at UCLA



Stacked Patch with Center-Feeding



Front-view of Stacked Patch with fixture



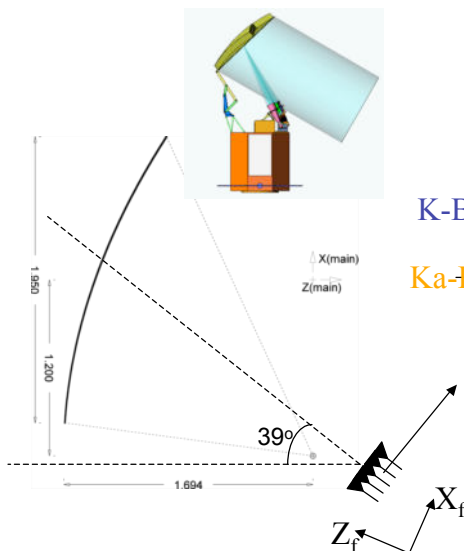
L-Band Probe used for Measurement



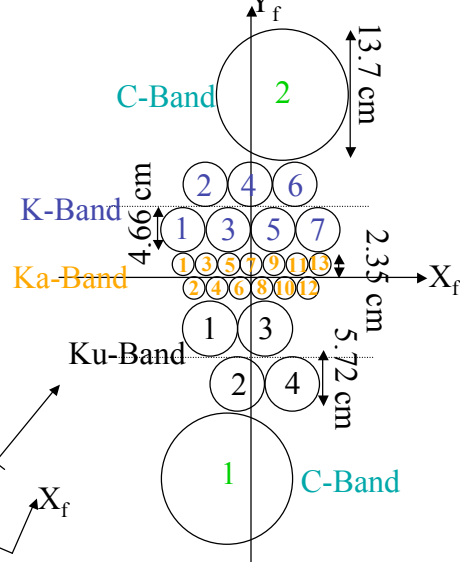
Near-Field Measurement Chamber

An Offset Parabolic Reflector Antenna & its Feed Farm

Antenna Geometry



Feed Configuration

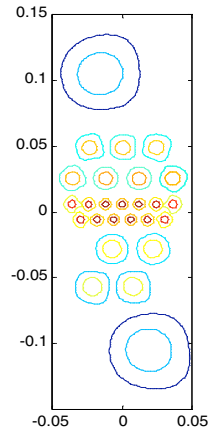


Note: In this figure Y is the flight direction.

An Offset Parabolic Reflector:

The overall beam footprints

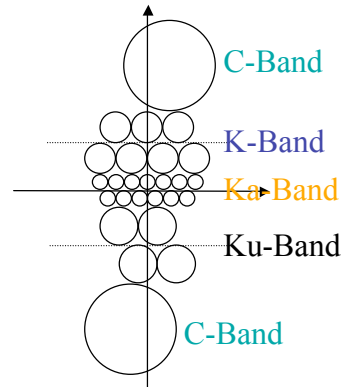
Far-Field Footprints



Inner Contour= -3 dB

2nd Contour= -10 dB

Feed Configuration

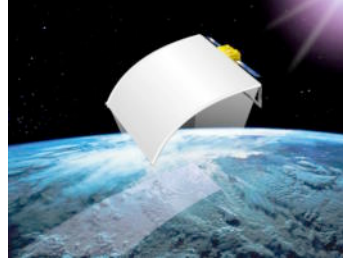


**5.3 m by 5.3 m Offset Cylindrical Parabolic
Reflector Antenna with Linear Array Feed for
1-D Beam Steering at Ku and Ka Bands**

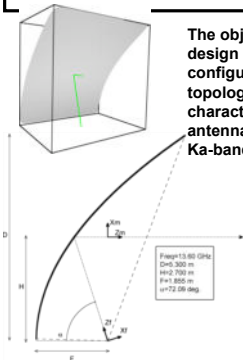
Precipitation Radar Antenna (PR-2) : Offset Parabolic Cylindrical Reflector Antenna

Novel Features:

- Incorporates dual-frequency with matched beams.
- Improves horizontal resolution to achieve wide-angle scanning of 30 deg.
- Avoids feed blockage.
- Better cross-polarization isolation as compared to a parabolic reflector.
- Light weight.

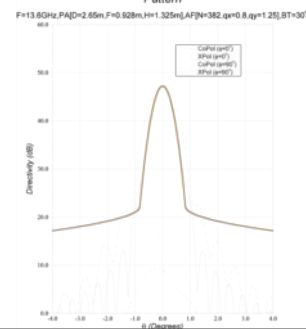
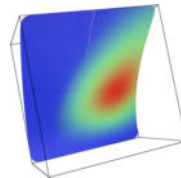


5.3 meter antenna concept Platform

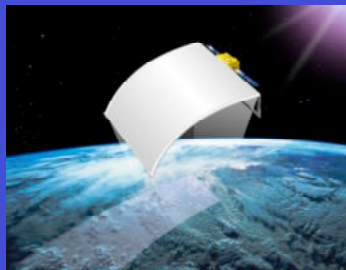


The objectives are: (i) the optimal design of the antenna configuration, (ii) feed array topology selection and (iii) characterization of the overall antenna performance at Ku and Ka-band frequencies.

Current



14/35-GHz Doppler Precipitation Radar



Current Status

IIP PR-2 and APRA Tasks:

- 2.7mx2.7m Ku- and Ka-band
- 0 and 30° beam pointing
- Prototype model complete in 9/04



APRA Prototype: reflector & Ku-band feed

Tasks Needed for 6mx6m cross-track scanning ($\pm 35^\circ$) reflector

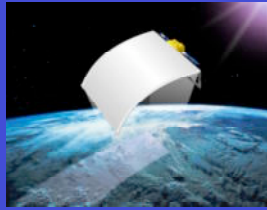
1. Space rigidization reflector structure
2. Longevity of membrane material for space
3. Ka-band T/R modules each at 0.75-W power level
4. Ka-band phase shifters with low insertion loss
5. Ka-band linear feed technologies (dual-polarization with 30-dB cross-pol isolation)
6. Verification method to compensate for gravitational loading (artificial) effect during ground testing
7. Metrology to detect the reflector surface distortion
8. Compensation of distorted reflector surface

Current TRL: 3
IIP Exit TRL: 5

Future Effort for reaching TRL=7

- Two years
- \$3M

14/35/94-GHz Doppler Precipitation/Cloud Radar



Current Status

IIP PR-2 and APRA Tasks:

- 2.7mx2.7m, 0/30° pointing Ku- and Ka-band
 - Prototype model complete in 9/04
- ESSP CloudSat radar:
- 2mx2m, nadir pointing W-band
 - Flight model completed in 3/03

Tasks Needed for 6mx6m cross-track scanning ($\pm 35^\circ$) reflector

1. Space rigidization reflector structure
2. Longevity of membrane material for space
3. Ka/W-band T/R modules each at 0.75-W power level
4. Ka/W-band phase shifters with low insertion loss
5. Ka/W-band linear feed technologies (dual-polarization with 30-dB cross-pol isolation)
6. Verification method to compensate for gravitational loading (artificial) effect during ground testing
7. Metrology to detect the reflector surface distortion
8. Compensation of distorted reflector surface

Current TRL: 1
IIP Exit TRL: 3

Future Effort for reaching TRL=7

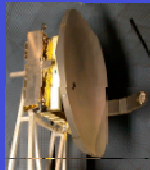
- Five years
- \$8M



APRA reflector



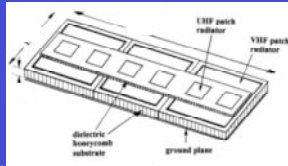
APRA Ku-band feed



CloudSat antenna

Array-Fed 30 m Parabolic Reflector Antenna Creating Matched Beams at VHF and P Bands

Multifrequency Dual-pol Feeds for Custom Aperture Formation on Large Reflectors

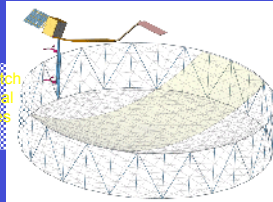


Current Status

MOSS: Dual-frequency double-stack planar feed will be at TRL-5 by end of 2004:

- 1:10 scale built/tested, currently being integrated with reflector for overall system testing
- full scale feed under construction
- achieves effective dual rectangular apertures on circular reflector

Current MOSS patch feed achieving dual rectangular apertures on MOSS



Tasks Needed for Multifrequency Dual-polarization Low-profile Feeds:

1. Relevant (launch environment) testing for MOSS' dual-frequency feed
2. Concurrent radar and radiometer system-level designs at multiple relevant frequencies (start with three)
3. Multiple stack planar feed design, including feed placement and movements
4. numerical simulation to demonstrate performance
5. Design of multifrequency power-dividing networks
6. Highest priority frequencies are 0.13-1.3 GHz
7. Fabrication and ground testing at 1:10 scale frequencies, then at full-scale
8. Testing of full scale engineering model at relevant launch environment

Current TRL (2 freqs): 5

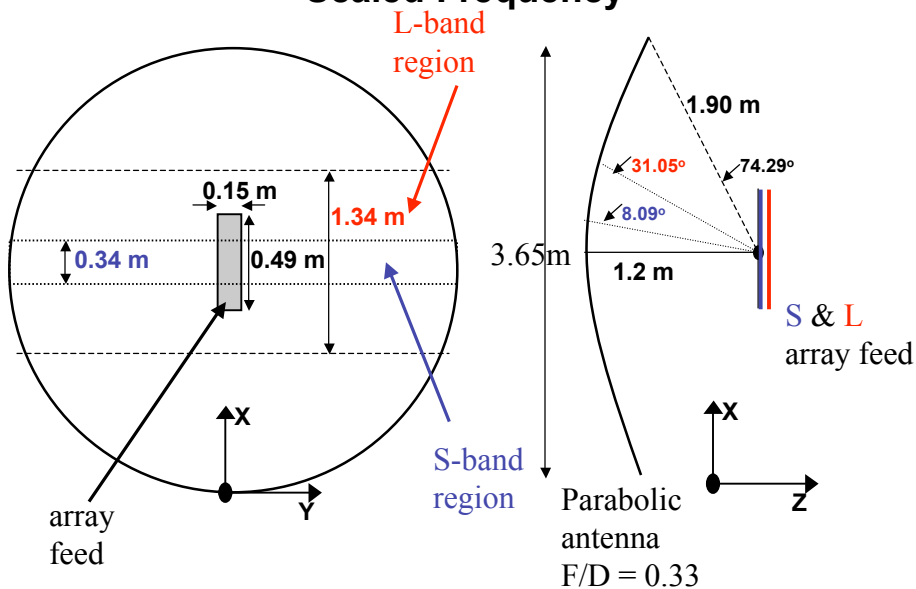
Current TRL (3 or more freqs): 2

Exit TRL: 6

Radar/Radiometer Working Group: 02/2004

3.65-m Reflector Antenna Geometry:

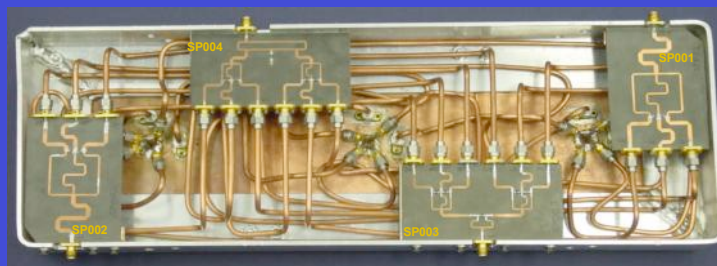
Scaled Frequency



Dual-Frequency Dual-Polarized Stacked Patch Microstrip Array

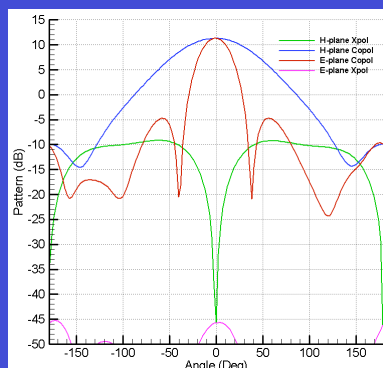


Front-View of the Array

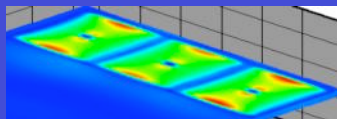


Back -View of the Array

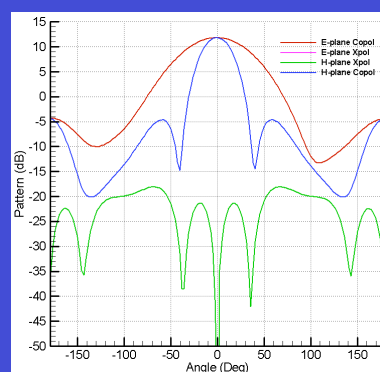
Simulation Results for the Array (Lower Patch)



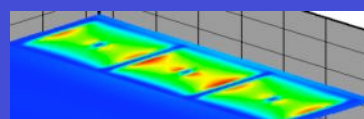
Radiation Pattern (Array Polarized)



Simulated Current Distribution on Array (at last time-step)
(Array Polarized)



Radiation Pattern (Cross-Array Polarized)

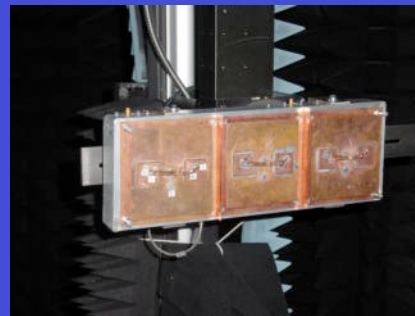
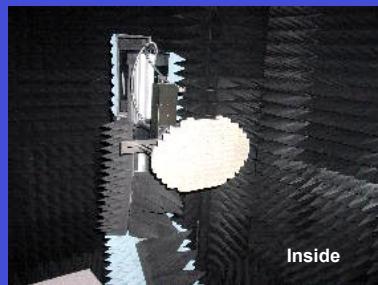


Simulated Current Distribution on Array (at last time-step)
(Cross-Array Polarized)

UCLA Spherical Near Field Measurement Chamber



Array Antenna in the Chamber

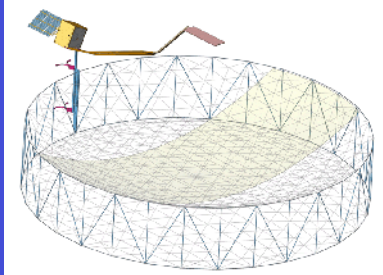


Parabolic Reflector to be Fed by the Array: Proof of Concept



3.65-m parabolic
reflector antenna
with $F/D = 0.33$

30m AM2 Mesh Reflector, 0.13-1.3 GHz



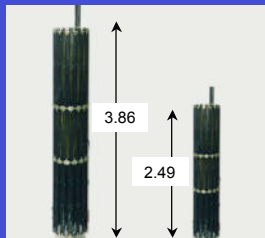
AM2 30-m reflector+feed and S/C point design for MOSS

Tasks Needed for 12-50m AM2 reflector

- (2) scaled engineering model (12m) of AM-2 reflector technology
- (3) Component level ground verification of all "new" full-scale components and subsystems in their relevant environment(s)
- (4) Testing a scaled AM2 antenna system (such as existing 12-meter engineering model) in the relevant launch environment
- (5) Establish the remainder of TRL 6 status by similarity to heritage
- (6) Feed technologies (for example, planar and multi-frequency feeds - see associated slide)

Current Status

1. Thuraya: 12m diameter at TRL 9, with Astro-Mesh-1 (AM1) technology
2. Northrop-Grumman (Astro) will take the new Astro-Mesh-2 (AM2) technology to TRL 5 in near future. AM2 has considerably lower package volume compared to AM1, and is applicable to up to 50-m class reflectors. TRL 6 is achievable in 2006-2007 timeframe.

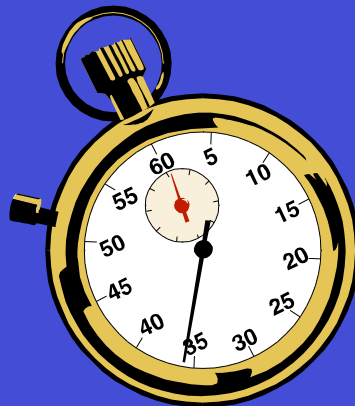


Package height in meters: AM1 (left) and AM2 (right) for a 12m reflector

Current TRL: 5
Exit TRL: 6

Radant/Radiometer Working Group, 02/2004

How am I doing with the time?

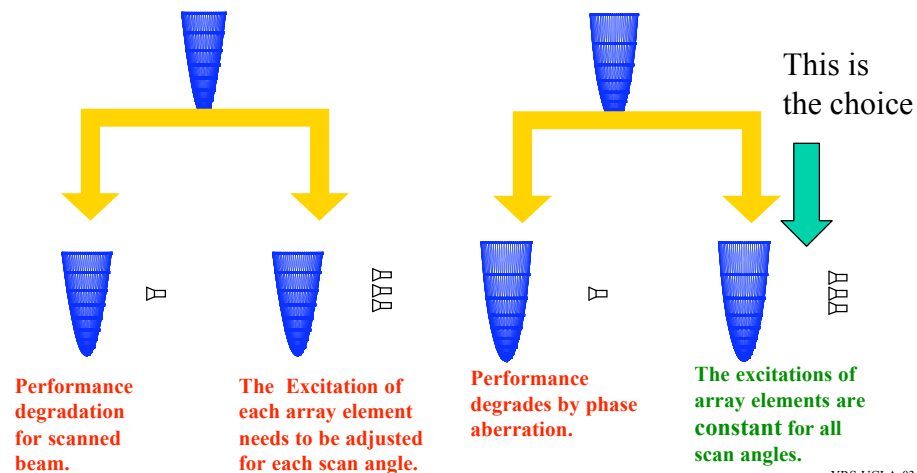


Array Compensated 35 m Spherical Reflector Antenna with Spiral Feed Motion for Producing Large Scan at Ka Band

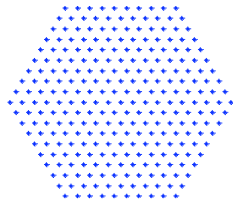
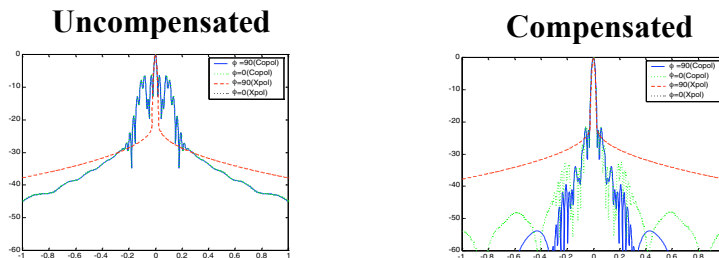
Potential Design Concepts

Parabolic Reflector

Spherical Reflector



Array Compensated Spherical Reflector Antenna Freq = 35.6 GHz



	Un compensated	Compensated
Directivity (dB)	67.62	77.48
HPBW (deg.)	0.018	0.022

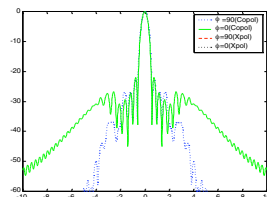
Noticeable improvement is observed.

YRS-UCLA-03

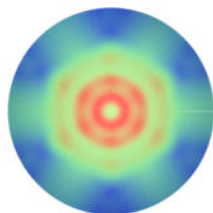
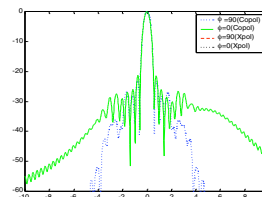
1.5 meter Spherical Reflector Antenna Scanning Performance

$F/D = 0.4$, $F/\text{Dill} = 0.45$
Hexagonal Array

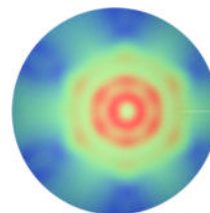
No Scan Case
Directivity = 51.47 dB



4° Scan Case
Directivity = 51.42 dB

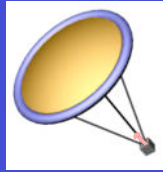


**Surface
Current on the
reflector**

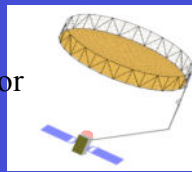


YRS-UCLA-03

35-GHz Geostationary Doppler Hurricane Radar



or



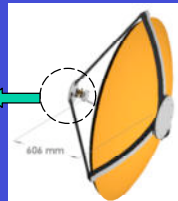
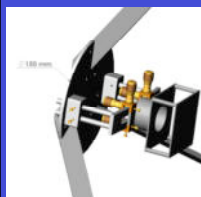
Tasks Needed for 35m spiral-scan reflector

1. Light weight, rigid, deployable spherical reflector
2. Longevity of membrane/mesh material for space
3. Innovative low-loss power dividing network Ka-band low-loss planar array feed technologies
4. Innovative electro-mechanism for spiral scanning for twin array feeds
5. Ground testing and verification method (hybrid of measurements and simulation)
6. Verification method to compensate for gravitational loading (artificial) effect during ground testing
7. Metrology to detect the reflector surface distortion
8. Adaptive compensation of distorted reflector surface

Current Status

IIP Nexrad-In-Space (NIS) Tasks:

- 1.5m (dia.) Ka-band variable pointing prototype
- Electro-mechanical scanner demonstration model
- Models to be completed in 10/05



NIS Prototype Antenna Design

Current TRL: 2

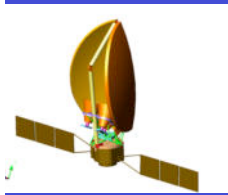
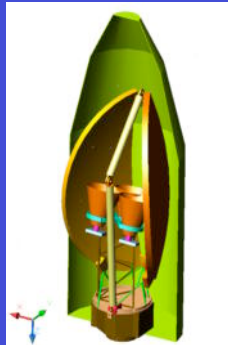
IIP Exit TRL: 4

Future Effort for reaching TRL=7

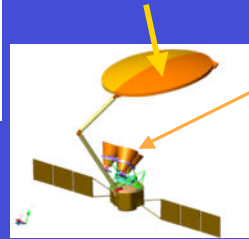
- Five years (from IIP Exit in 10/05)
- \$6M (exclude 35-m reflector)
 - Leverage off on-going reflector development

Spinning (and non spinning) Reflector Antenna Designs at L Bands

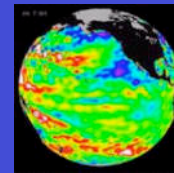
Aquarius: Global maps of ocean-salt concentration and Soil Moisture



3 m foldable
reflector antenna for
earth observation

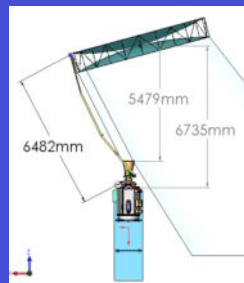


GA
Optimized
horns



E. G. Njoku, W. J. Wilson, S. H. Yueh and Y. Rahmat-Samii, "A Large-Antenna Microwave Radiometer-Scatterometer Concept for Ocean Salinity and Soil Moisture Sensing," *IEEE Transactions on Geoscience and Remote Sensing*, Vol. 38, No. 6, pp. 2645-2655, Nov., 2000.

Large Spinning Reflector



Tasks Needed for 6-30m rotating reflector

1. Light weight, rigid, deployable reflector (highest frequency Ku-band)
2. Finite element analysis tools
3. Ground testing and verification method
4. Compensation of distorted reflector surface
5. Metrology to detect the reflector surface distortion

Current Status

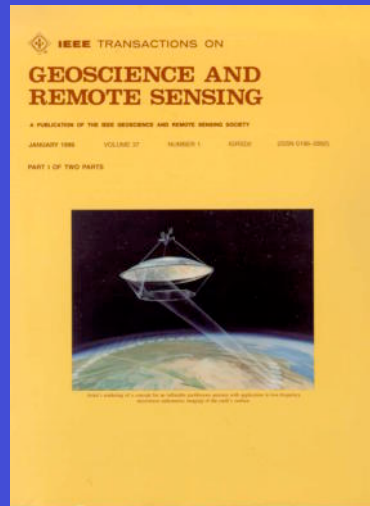
1. HYDROS: 6m (diameter) rotating antenna at L-band
2. HYDROS will improve TRL 4 to TRL 6

Current TRL: 3
Exit TRL: 6



Example: NGST Astro deployable mesh antenna

Torus-class Antennas can provide viable solutions for overcoming spinning large reflector antennas



**There are many varied
configurations for
torus class reflector
antennas which
require further studies.**

Large Membrane Array and Reflectarray Antennas at L, X and Ka Bands

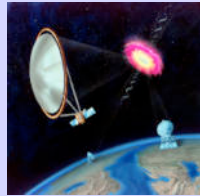
How to obtain LARGE apertures?

Filled apertures:

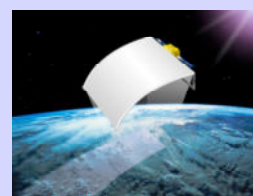
2-D curved



2-D curved

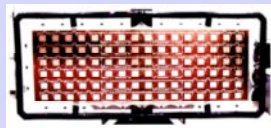


1-D Curved



Discrete dense apertures:

Every Element is Fed

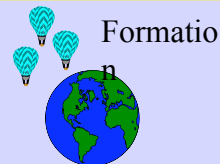
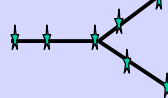


Space Fed



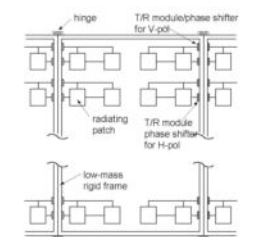
Discrete sparse apertures:

Sticks

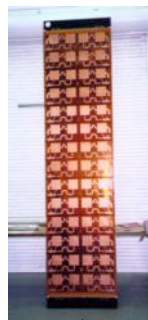


Microstrip Array: Discrete Dense Aperture Element Fed

Foldable Framed Thin-Membrane L-Band Array



Mass density = 1.8 kg/m²



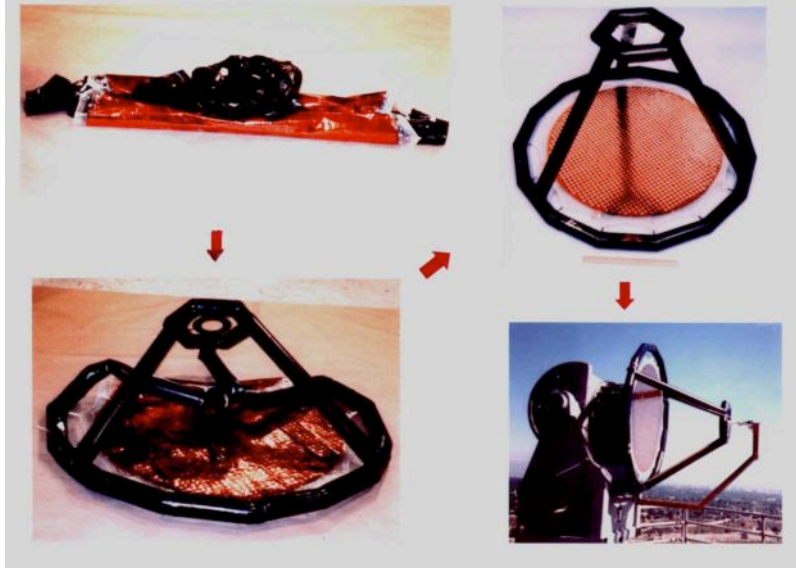
Each panel



Close-up view
2-layer thin-membrane

Full-size 5m x 3m 7-panel array

JPL 1 m X-Band Inflatable Microstrip Reflectarray



Reflectarray: Discrete dense aperture with space-fed elements

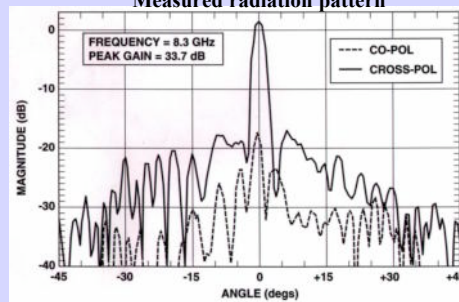
1m X-band reflectarray



By JPL / ILC Dover

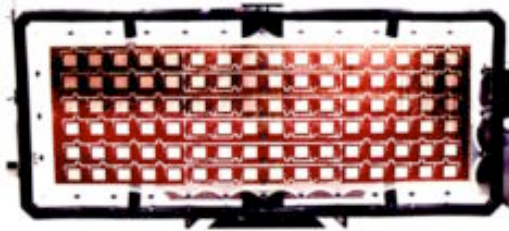
- Circularly polarized
- Approx. 1000 elements
- 2-layer membranes
- Patch/ground plane separation is 1.3mm
- Kevlar tube / Kapton membrane
- Radiation efficiency = 37%
- Achieved surface flatness < 0.8mm

Measured radiation pattern



SAR Arrays: Discrete Dense Aperture Element Fed

Two L-band inflatable SAR arrays were developed



ILC Dover design

L'Garde design

- Antenna concepts and RF designs were done by JPL, inflatable structures were done by ILC Dover and L'Garde
- Both models use inflatable tubes to support a multilayer membrane aperture
- Material: Kevlar tube/Kapton membrane
- Efficiency achieved: ILC Dover 52%
L'Garde 75%
- Achieved surface flatness < 0.8mm



Concluding Remarks

Observations

Existing programs have advanced our understanding of potential applications of large deployable antennas.

Follow up research and development activities are needed to make them technologically ready for spaceborne utilization.

Technological readiness for Reflector Antennas: Membrane/Inflatable Reflector Antennas

No existing commercial heritage for offset antennas.

Utilization on a spinning platform requires evaluation.

Surface accuracy and profile needs characterization.

Utilization of an adaptive surface correcting system will enhance performance at high frequencies.

Some of the concepts could be evolved to 25-m aperture dimensions.

Both ground and in-flight characterizations/diagnostics are necessary.

Technological readiness for Reflector Antennas: **Mesh Deployable Reflector Antennas**

Existing commercial heritage for up to 12.5m offset antennas.

Utilization on a spinning platform requires evaluation.

Needs to be evaluated for frequencies beyond 20 GHz.

Surface accuracy and profile needs characterization and adaptive correction may be necessary.

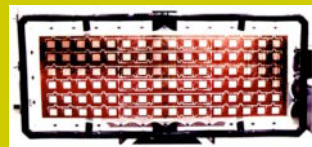
Some of the concepts could be evolved to 25-m aperture dimensions.

In-flight demonstration will create confidence.

The Development of Inflatable Arrays

Future challenges in the RF area:

- **Bandwidth improvement > 15%**
- **Dual-band shared aperture**
- **Membrane mounted T/R module/phase shifter with dc control circuitry**
- **RF power distribution on large aperture (> 10m)**
- **Noise temperature & beam efficiency characterization**





Thank you